

Fitting GRB211211A with a photospheric model including radiation-mediated shocks

Oscar Wistemar¹ Filip Alamaa^{1,2}, Felix Ryde¹



wistemar@kth.se

Photospheric spectrum from GRB211211A

A photospheric model including radiation-mediated shocks (RMSs) is used to fit data from GRB211211A. The model fits very well, even better than the typically used Band function. This shows that photospheric spectra can be both soft and broad enough to describe typical GRB observations.

Constraining physical parameters

Within the framework of the photospheric model we can derive the physical RMS parameters. Examplifying with time bin 19 (7.2-7.3 s), we find that the data is well described by a strong sub-photospheric shock (high $\gamma_u \beta_u$) in a hot outflow (high θ_u) at a moderate optical depth ($\tau \sim 20$) in a photon rich environment ($n_{\gamma}/n_p \sim 3 \times 10^5$).

Please ask me about the figures!

The Kompaneets RMS Approximation (KRA)

This is made possible thanks to the Kompaneets RMS approximation (KRA; Samuelsson et al. 2022), which is an accurate and extremely fast way to model sub-photospheric dissipation through radiation-mediated shocks in GRBs. The low computational cost of the model allows us to study these types of shocks in a quantitative way.



Example fit at 7.2 - 7.3 s

Corner plot of KRA fit at 7.2 – 7.3 s





Affiliations

KTH Royal Institute of Technology, Stockholm, Sweden 2 IAP Institut d'Astrophysique de Paris

Further Reading





Samuelsson, Lundman, and Ryde (2022)

Samuelsson and Ryde (2023)