Synchrotron polarization of GRB afterglow shocks with hydrodynamic-scale turbulent magnetic field

Kuwata et al. 2023, ApJ, 943, 118, Kuwata et al. in prep

Asuka KUWATA, K.Toma, S.S.Kimura, S.Tomita (Tohoku U., Japan), J.Shimoda (Tokyo U.)



Abstract

We construct a semi-analytic model of afterglow polarization with hydrodynamic-scale turbulent B-field. We find that for the isotropic turbulence and the zero-viewing angle, $\Pi_{radio} \sim \Pi_{opt} \sim 2f_B \%$ on average but Π_{radio} can be higher than Π_{opt} at some time intervals. PDs & PAs vary randomly and continuously at both of two bands. On the other hand, for the anisotropic turbulence and the finite viewing angle, the optical polarization shows similar behavior to the plasma-scale field model around the jet break time. These results suggest that hydrodynamic-scale model is consistent with all polarimetric observational data of afterglows.

1. What is the origin of the strong magnetic field?





Summarv hydrodynamic-scale model is consistent with obs.